



1  
00:00:02,000 --> 00:00:05,290  
[Man off screen] Try this one on for size,  
our sun's north and south poles are about

2  
00:00:05,290 --> 00:00:06,460  
to flip flop.

3  
00:00:06,460 --> 00:00:09,920  
That's right if you were near the Sun and  
had your compass pointing north the dial would

4  
00:00:09,920 --> 00:00:11,870  
actually swing south.

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00:00:11,870 --> 00:00:15,960  
Joining us today from the Goddard Space Flight  
Center in Greenbelt Maryland is NASA Heliophysicist

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00:00:15,960 --> 00:00:18,109  
Dr. Alex Young, thanks for joining us.

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00:00:18,109 --> 00:00:19,109  
Alex: Good morning.

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00:00:19,109 --> 00:00:21,180  
[Man off screen] What is a magnetic field?

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00:00:21,180 --> 00:00:22,490  
And why do we study them?

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00:00:22,490 --> 00:00:28,820  
Alex: Well when you stick a magnet to a refrigerator,  
that happens because of this invisible field

11  
00:00:28,820 --> 00:00:30,360  
called a magnetic field.

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00:00:30,360 --> 00:00:37,130

The sun has a magnetic field also and the Sun's magnetic field is incredibly complicated

13  
00:00:37,130 --> 00:00:43,600  
and dynamic and it drives all of the activity that happens on the sun, space weather and

14  
00:00:43,600 --> 00:00:45,050  
solar storms.

15  
00:00:45,050 --> 00:00:50,320  
So we want to understand these magnetic fields so that we can understand these storms why

16  
00:00:50,320 --> 00:00:54,820  
they happen and how powerful they could be.

17  
00:00:54,820 --> 00:00:59,140  
[Man off screen] What causes the Sun's magnetic field to flip?

18  
00:00:59,140 --> 00:01:05,039  
Alex: Well inside the sun, it's not a solid body, so it rotates at different speeds at

19  
00:01:05,039 --> 00:01:12,080  
different places and this generates magnetic fields via something called a "solar dynamo".

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00:01:12,080 --> 00:01:17,270  
These magnetic fields are generated and they increase in strength rising to the surface

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00:01:17,270 --> 00:01:23,290  
giving us sun spots, and this gives us the solar activity that we're talking about.

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00:01:23,290 --> 00:01:30,900  
This whole process happens over a time scale of roughly 11 years we call this the "solar

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00:01:30,900 --> 00:01:32,060  
cycle".

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00:01:32,060 --> 00:01:39,270  
The solar cycle goes from being very weak  
with low activity to high activity and then

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00:01:39,270 --> 00:01:40,810  
back down to low activity.

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00:01:40,810 --> 00:01:47,909  
But what's exciting for us right now is that  
we are near this peak of activity and what

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00:01:47,909 --> 00:01:57,000  
also happens with that is that the orientation  
of the Sun's magnetic field reverses itself,

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00:01:57,000 --> 00:01:58,360  
switches places.

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00:01:58,360 --> 00:02:04,560  
And so this is what's happening right now  
because we are near this solar maximum period.

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00:02:04,560 --> 00:02:08,310  
[Man off screen] What does this mean for the  
Earth?

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00:02:08,310 --> 00:02:16,360  
Alex: Well, all of this activity, solar flares  
(these bright flashes of light) or coronal

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00:02:16,360 --> 00:02:20,370  
mass ejections which are huge explosions of  
material.

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00:02:20,370 --> 00:02:25,230

They don't impact us directly but what they do impact is our technology.

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00:02:25,230 --> 00:02:31,150

They can cause temporary disruptions in our communications, GPS, they can create a dangerous

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00:02:31,150 --> 00:02:37,470

environment for astronauts in space, and they can also cause disruptions in our power grid.

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00:02:37,470 --> 00:02:43,510

Now the one good side about that is it also gives us the aurora that we can see at the

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00:02:43,510 --> 00:02:44,879

northern and southern poles.

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00:02:44,879 --> 00:02:49,180

[Man off screen] Well where can we see more great images of the sun and learn about the

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00:02:49,180 --> 00:02:50,200

current solar cycle?

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00:02:50,200 --> 00:02:54,940

Alex: Well a great place to go to find out all about this and see all of the great data

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00:02:54,940 --> 00:02:59,099

we have and imagery is at [www.nasa.gov/sunearth](http://www.nasa.gov/sunearth).